

Dynamics of the Mars Atmosphere: The Post-Viking Perspective

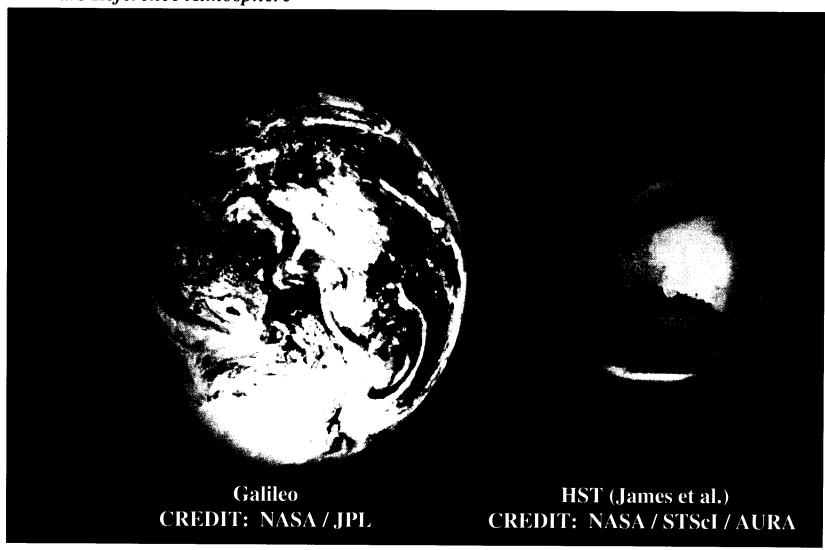
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*Work performed at the Jet Propulsion Laboratory, California Institute of Technology, under contract with the National Aeronautics and Space Administration

Earth & Mars



— Mars Reference Atmosphere



1982 COSPAR Mars Reference Atmosphere (1)



- Mars Reference Atmosphere
- Post-Viking Models of Summer Atmospheric Structure
 - 2 Viking Entry Profiles
 - Seasonal Variation of Surface Pressure
- Martian Meteorological Variability
 - Seasonal wave phenomena in lower atmosphere
 - Diurnal and semidiurnal atmospheric tides (daily variations in V,T,P)
 - Surface wind measurements at the two Viking Lander sites
 - Evidence in cross-equatorial Hadley circulation in T, sfc. P
 - Heavy reliance on models to interpret data
- Properties and Impact of Dust in the Martian Atmosphere
- IRTM: Temperature and Dust from IR data
 - Wintertime Polar Warmings
 - Dust opacities from thermal IR data
 - Diurnal ground temperature variations

1982 COSPAR Mars Reference Atmosphere (2)



- Mars Reference Atmosphere
- Clouds of Mars
 - Dust hazes, local dust storms and major dust storms
 - Condensate clouds: orographically forced, polar hood, storms
 - Surface frosts (snow?)
- Composition
- Water Vapor (column abundances for most of a Mars yr.)
- Mars Upper Atmosphere: Mean and Variations
 - Dust Storm effects
 - Diurnal variations
 - Solar Cycle Variations (Mariners 6, 7, 9, and Viking)
- Upper Atmosphere Wave Structure
 - High altitude wave structure (in temperature profiles)

1992 MARS Book: Issues in Atmos. Dynamics



— Mars Reference Atmosphere

Hadley-type Circulations

- Strength, Extent, Nature (e.g., 1 cell or 2 at equinox; vertical range)

Atmospheric Waves

- Strength of wave activity at high southern latitudes during winter
- Influence of topography (orography and variations in surface A,I)
- Influence of gravity waves and nature of polar warmings

Dust Cycle

- The nature of great dust storms and their interannual variation
- Feedback mechanisms

Role of Transport in Water and Carbon Dioxide Cycles

- Transport of mass and of heat into and/or out of the polar regions

Improvements in Atmospheric Modeling

- Increased spatial resolution and vertical range
- Better parameterizations of dust movement, injection and removal

Atmospheric Data Sets after Viking (1 of 2)



- Mars Reference Atmosphere
- Earth-Based Observations (continuing)
 - Ground-based observations at microwave wavelengths
 - Hubble Space Telescope
- Phobos-2 Mission (Feb.-March, 1989)
 - Termoscan: surface temperatures and atmospheric dust properties
 - Solar Occultations: atmospheric profiles (T, aerosols, condensates, water)
- Mars Pathfinder (July-Sept., 1997)
 - IMP: Sky views
 - Atmospheric opacity, cloud and haze structure
 - Mesoscale phenomena (e.g., dust devils)
 - MET: Near-surface pressure, temperature and wind
 - Diurnal variations
 - Dust Devil signatures

Atmospheric Data Sets after Viking (2 of 2)



— Mars Reference Atmosphere

Mars Global Surveyor (1997 -)

- MOLA: Topography and mass balance
- TES: Atmospheric profiles, polar cap monitoring, surface properties
- MOC: Daily global mosaics of atmospheric & surface phenomena
- RS: High-vertical-resolution profiles from radio occultations
- MAG/ER: Upper atmospheric structure and magnetic field
- ACCEL: Upper atmospheric densities & scale heights during aerobraking

• 2001 Mars Odyssey (2001 -)

- THEMIS: Monitoring mid-level atmospheric T & polar phenomena
- GRS: Detection of water (hydrogen) in topmost meter of ground
- ACCEL: Upper atmospheric densities & scale heights during aerobraking

Major Advances since Viking (my view!)



- Mars Reference Atmosphere
- Atmospheric water is a major player in Mars climate through its role in removing atmospheric dust [Clancy, Rodin et al.]
 - Ice condensation and scavenging of dust is a major control of water vapor and atmospheric dust distributions in the current climate and potentially in past climates
 - Drives part of the difference between northern & southern hemispheric seasonal activity; probably not a driver of interannual variability
- Large-scale topographic differences exert a major control on atmospheric circulation and transport [Richardson, Wilson]
 - Western boundary currents, northern wave activity
 - Control on Hadley-type circulations may damp variations due to orbital variations on astronomical time scales

Major Advances since Viking (my view!)



— Mars Reference Atmosphere

(continued)

- Seasonal variations are very predictable up to a point [Malin et al.]
 - Onset of local dust storms & mesoscale activity (e.g., dust devils)
 - Once dust-raising systems are active (and ice-scavenging removal is lessened?), atmosphere is less predictable
 - \Rightarrow Northern spring and summer seasons are less variable (up to Ls ~ 165?) than southern spring and summer
- Repeating seasonal activity may trigger interannual variation [*Leovy et al.*] but triggers for southern subtropical dust storms may come from north (*Richardson et al.*) or south?!
- Waves generated in the lower atmosphere really do make into the upper atmosphere, particularly the diurnal Kelvin wave [Keating, Wilson, Forbes, et al.]

HST View in 2001



— Mars Reference Atmosphere

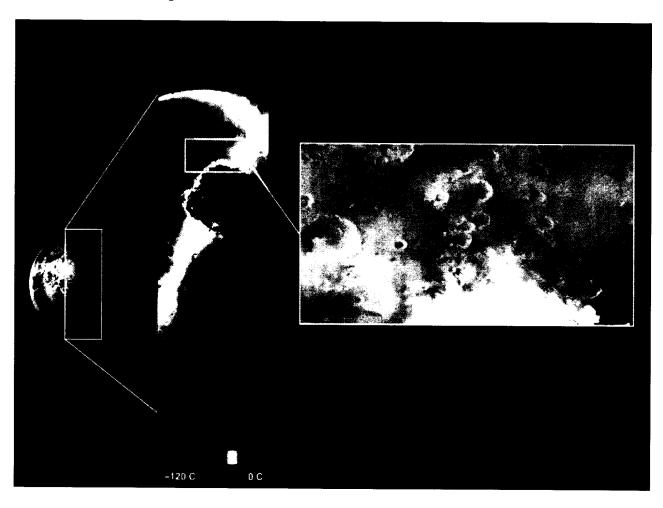


CREDIT: NASA/STScI/AURA/J. Bell & M. Wolff

South Pole in Late Summer



— Mars Reference Atmosphere



Impact of New Data Sets



— Mars Reference Atmosphere

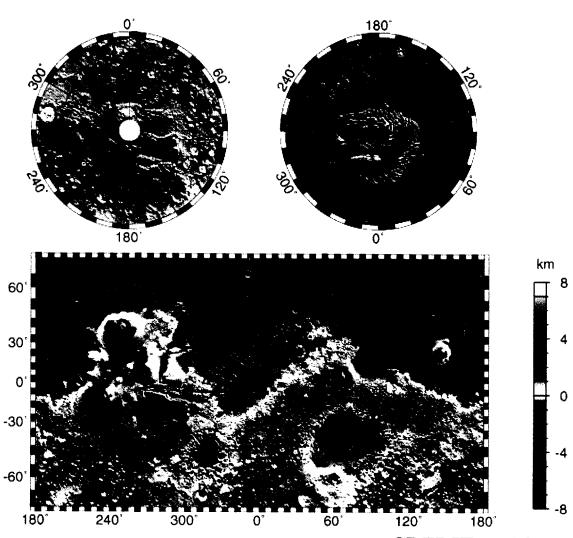
Of the new data sets, Mars Global Surveyor is likely to have the greatest impact on understanding atmospheric phenomena:

- Systematic Coverage
 - Daily, Global Sampling key to understanding time-varying fields
 - Multiple Observing Systems (e.g., visible and thermal IR)
- Long Record
 - Comparison of different Mars years is key to understanding Interannual Variability
 - MGS now in third Mars year of observations
 - MGS observational record includes years with and years without a global dust storm

Mars Topography (MGS MOLA)



— Mars Reference Atmosphere

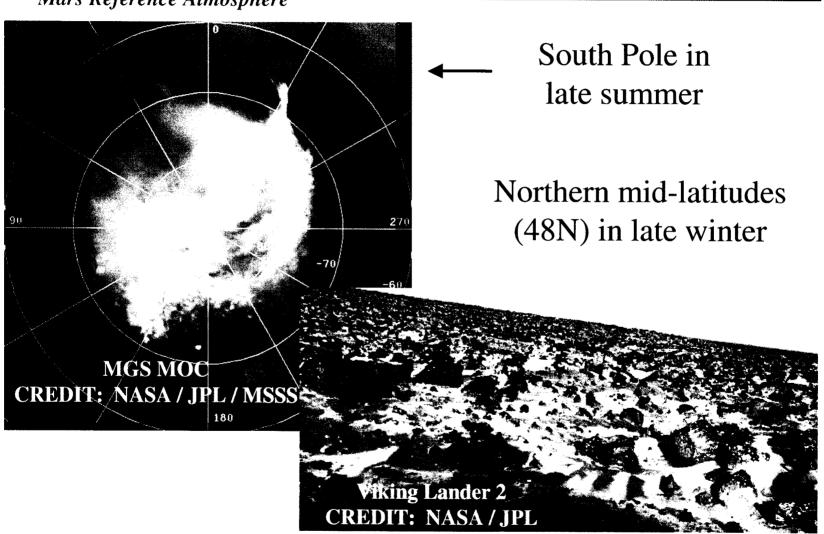


CREDIT: NASA/JPL/68FC

CO₂ Cycle: Seasonal Frosts



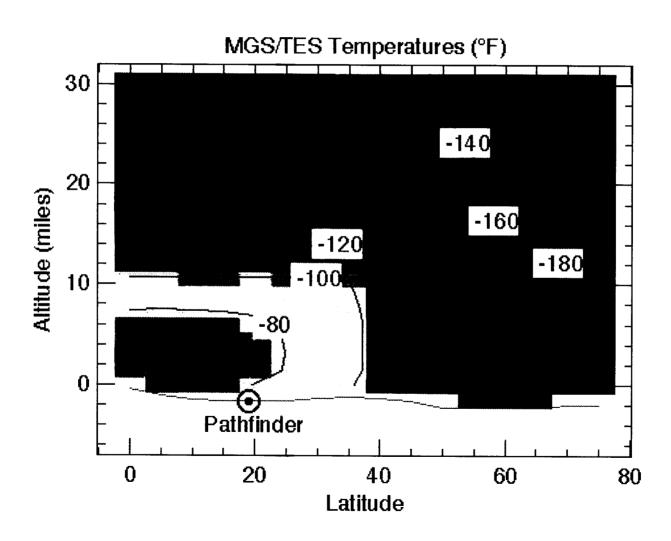
- Mars Reference Atmosphere



Zonal Temperatures in N. Winter



Mars Reference Atmosphere

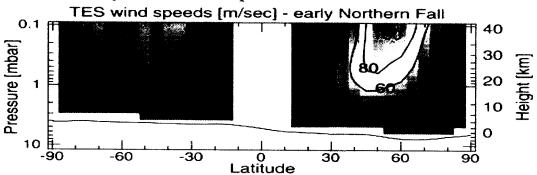


CREDIT: NASA/JPIrwzASU

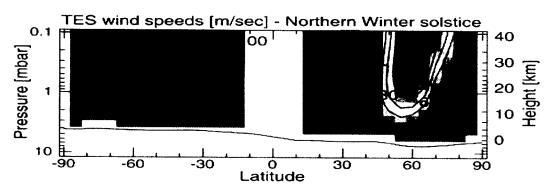
Geostrophic Winds



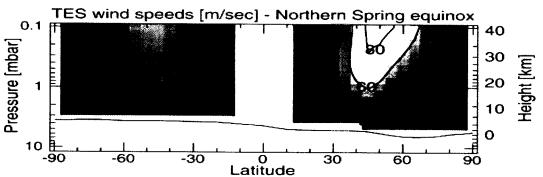




Early Southern Spring Early Northern Fall



N. Winter Solstice S. Summer Solstice (near perihelion)



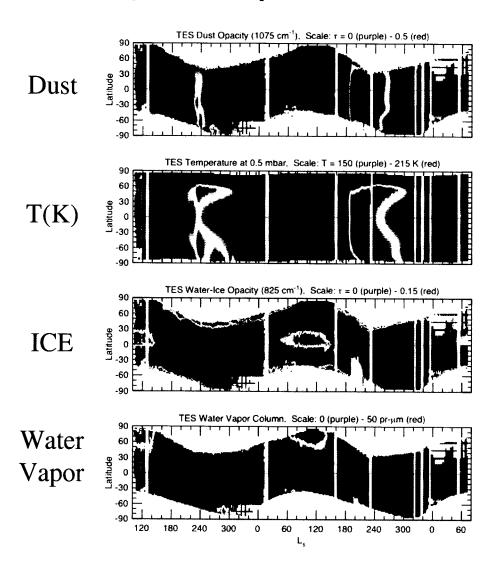
Vernal Equinox

CREDIT: NASA/JPL/ASU6

MGS TES Climatology



— Mars Reference Atmosphere



Zonally Averaged Fields during Two Mars Years

Courtesy of M. Smith TES Team

(TES PI: P. Christensen)
NASA/JPL/ASU/GSFC

Equatorial Cloud Belt (Mars ITCZ?)



— Mars Reference Atmosphere

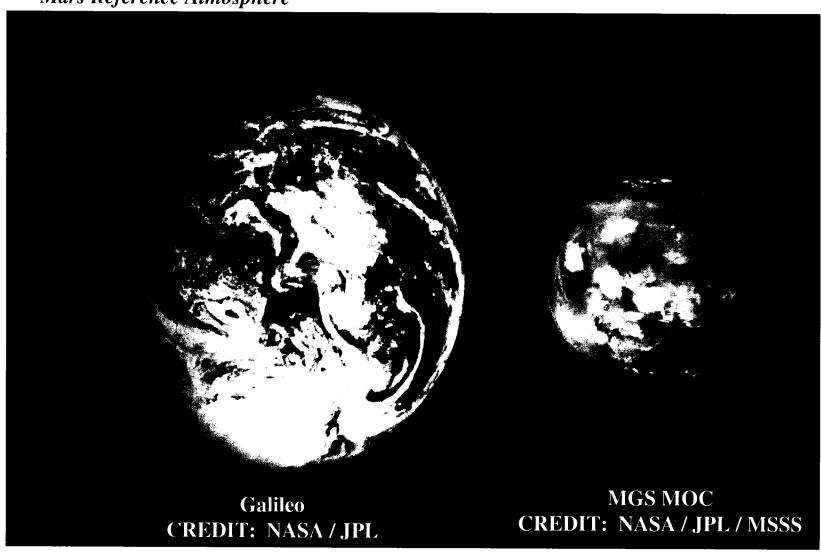


MGS MOC (PI: M. Malin)
CREDIT: NASA / JPL / MSSS

Earth & Mars



— Mars Reference Atmosphere



How to Raise Dust on Mars



Mars Reference Atmosphere

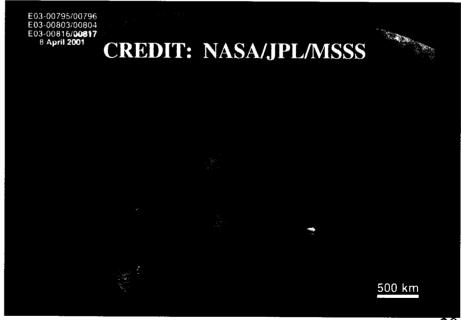


MGS Mars Orbiter Camera (MOC)

PI: M. Malin (MSSS)

Dust Devil Tracks

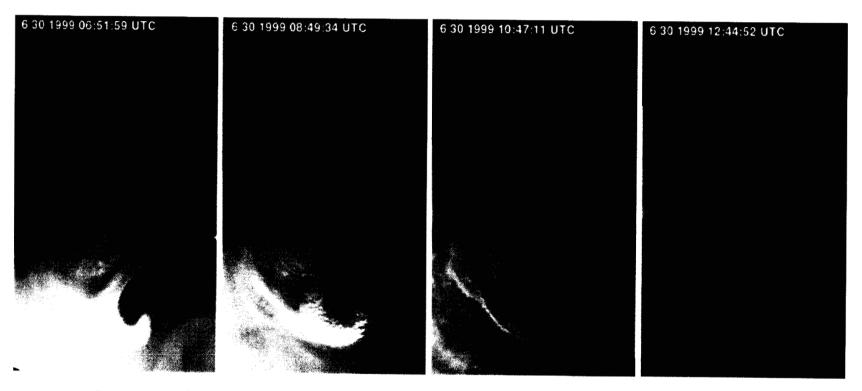
Local Dust Storm at South Pole Cap Edge



North Polar Dust Storm



— Mars Reference Atmosphere



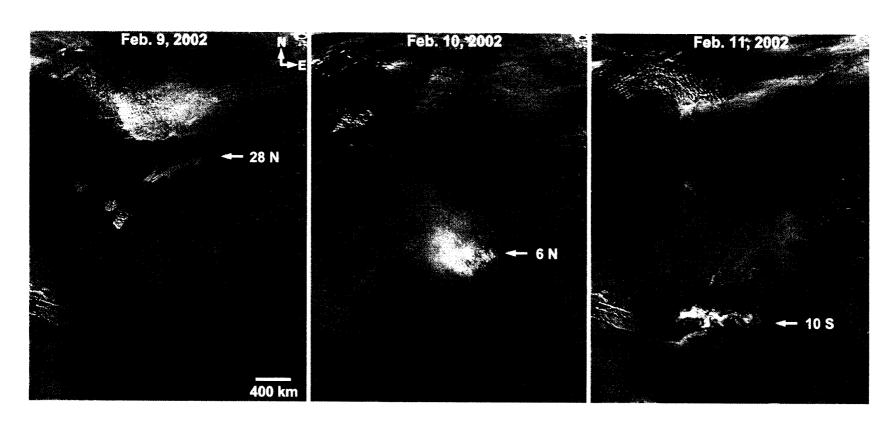
Viewed on Consecutive MGS Orbits (~ 2 hrs apart)

MGS MOC CREDIT: NASA / JPL / MSSS

Northern Storm Paths



— Mars Reference Atmosphere

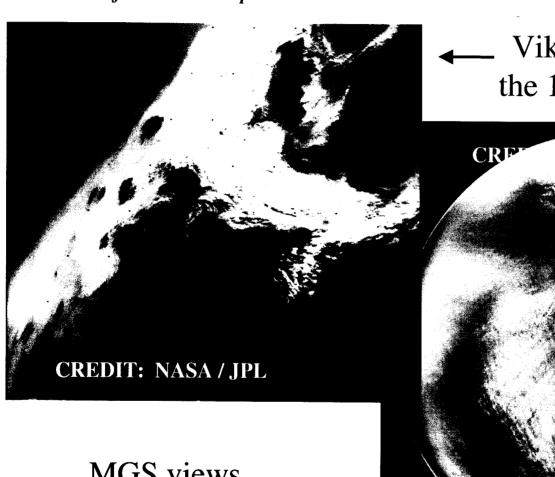


MGS MOC CREDIT: NASA / JPL / MSSS

Two Faces of Dust Storm Onset

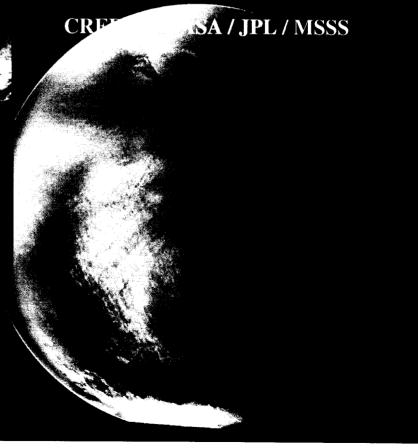


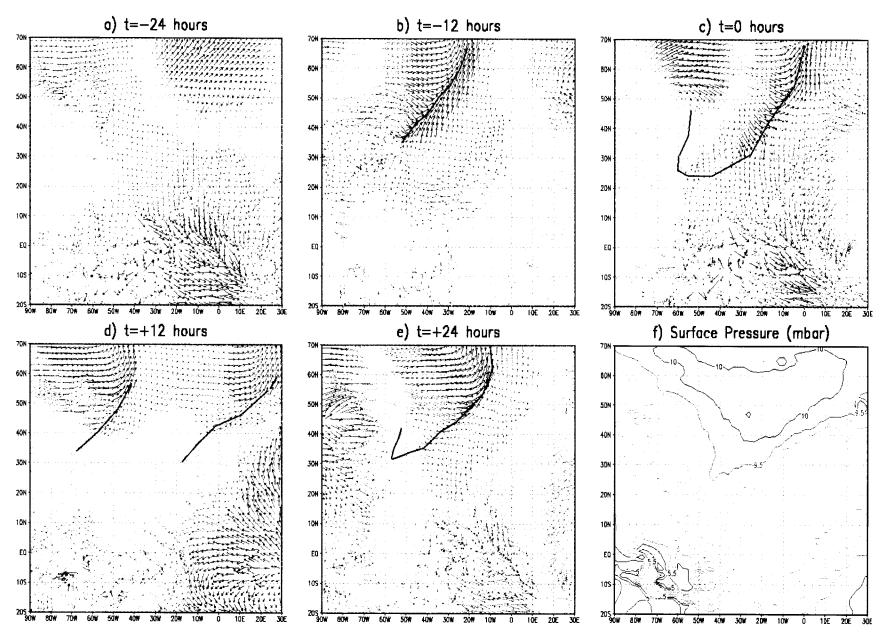
— Mars Reference Atmosphere



MGS views __

Viking views
 the 1977a GDS



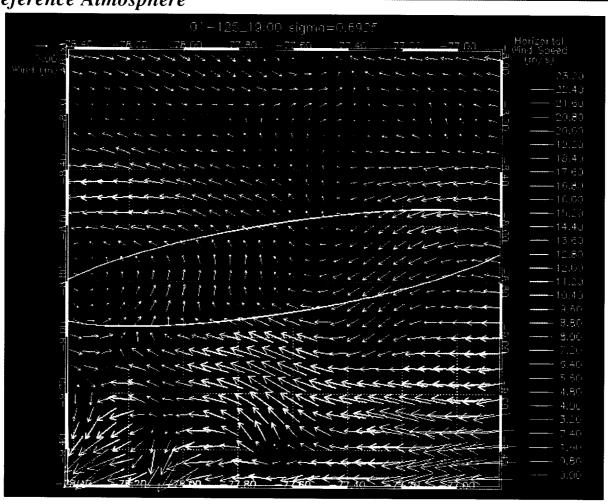


TIDAL GATE for FRONTS: H. Wang, et al., manuscript in preparation rwz-24

Mesoscale Modeling: Melas Chasma



— Mars Reference Atmosphere



Courtesy of S. Rafkin: MRAMS Simulation for MER

Limitations of Current Data



- Mars Reference Atmosphere
- Current Limitations w.r.t. Dynamical Meteorology
 - Restricted altitude range for global coverage (< 80 km)
 - Restricted <u>local time coverage</u>
 - <u>Limited spatial resolution</u> (horizontal and vertical)
 - Especially near surface
 - Lack of concurrent surface pressure data
 - No independent wind data!
 - Little independent validation (e.g., using different data types)
- ⇒ Continued Dependence on Atmospheric Models for Scientific Understanding
 - Fortunately, atmospheric models have improved
 - Improved resolution, vertical range, & physical parameterization
 - Still need vigorous validation
 - Models can reveal observational faults or inconsistencies
- ⇒ Continued Dependence on Parametric Engineering Models

Future Atmospheric Data Sets (1 of 2)



— Mars Reference Atmosphere

Mars Express (arrives in 2003/2004)

- Elliptical 8 hr orbit, 250 km periapsis
- Planetary Fourier Spectrometer (PFS): Atmospheric Profiles (nadir viewing) and surface observations
- OMEGA: Column abundances of atmospheric water vapor and aerosols in addition to surface composition
- SPICAM: UV solar/stellar occultations and UV emission for atmospheric structure and aerosol abundances and upper atmospheric composition; NIR for column water amounts
- HRSC: High-resolution stereo imaging of mesoscale phenomena
- MARSIS: Polar ice cap structure

• **NOZOMI** (arrives in 2003/2004)

- Mission geometry in revision
- Upper atmospheric data at different local times

Future Atmospheric Data Sets (2 of 2)



— Mars Reference Atmosphere

Mars Reconnaissance Orbiter (arrives in 2006)

- Near-circular polar orbit, 265-320 km, frozen over S. Pole
- Mars Climate Sounder (MCS: PMIRR rebuild): Daily Limb Sounding and Polar Monitoring
- CRISM: Seasonal Monitoring with EPF-type profiles, column abundances of dust, water vapor
- MARCI: Daily, Global Monitoring in UV and VIS channels
- HiRISE: High-resolution imaging of mesoscale phenomena, including dust storm onsets
- GRAVITY & RS: Tracking mass movement; occultation data (?)
- SHARAD: Near-surface water/water ice reservoirs?

NetLanders/Premiere (launched in 2007?)

- Surface meteorological measurements at 4 stations
- IR/Microwave sounding from Orbit (?)
- Global Weather Monitoring (?)

Reference Atmosphere



- Mars Reference Atmosphere
- Build on MGS data
 - Use MGS data to validate atmospheric models
 - Use models to fill in unobserved domains and to extend to different times (especially other times of day, solar cycle)
 - Use models to estimate unobserved, or loosely constrained, fields (e.g., near-surface winds)
- Update as Odyssey, Mars Express, NOZOMI, MRO, Netlander data become available
- Explore engineering sensitivities parametrically
 - Use available data as broad constraints
 - Use atmospheric models to estimate potential scenarios

Current observations of Mars have shown a fascinating range of atmospheric structure, variability and interaction which continues to challenge us